

1 **Phylogenetic and taxonomic revisions of Jurassic sea**
2 **stars support a delayed evolutionary origin of the**
3 **Asteriidae**

4
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23

24 Abstract

25

26 Background. The superorder Forcipulatacea **comprise** a major clade of sea stars with
27 approximately 400 extant species across three orders (Forcipulatida, Brisingida, Zorocallida).
28 Over the past century, the systematics of Forcipulatacea have undergone multiple revisions by
29 various authors, with some considering numerous families such as Asteroiidae, Zoroasteridae,
30 Pedicellasteridae, Pycnopodiidae, Labidiasteridae, and Neomorphasteridae, while others
31 recognized only two families (i.e., Asteroiidae and Zoroasteridae). Recent molecular analyses
32 have shown the artificial nature of some of these groupings. Notably, four well-supported clades
33 (Zorocallida, Brisingida, Stichasteridae, and Asteroiidae) emerged from a synthesis of
34 morphological and molecular evidence. The majority of extinct forcipulatacean species have
35 been placed in modern families. However, many of these fossil species are in need of revision,
36 especially those species placed within the Asteroiidae, the largest of all forcipulatacean families.

37

38 Methods. In light of recent advancements in forcipulatacean systematics, we comprehensively
39 reassess six well-preserved Jurassic forcipulatacean taxa, including the earliest crown-group
40 members from the Hettangian **stage** (~201.4 Ma), and also describe two new Jurassic genera,
41 *Forbesasterias* gen. nov. and *Marbleaster* gen. nov. We assembled the largest and most
42 comprehensive phylogenetic matrix for this group, sampling 42 fossil and extant forcipulatacean
43 species for 120 morphological characters. To infer phylogenetic relationships and construct an
44 evolutionary timeline for the diversification of major clades, we conducted a Bayesian tip-dating
45 analysis incorporating the fossilized birth-death process. A total of 13 fossil species were

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46 sampled in our analysis, including six taxonomically revaluated herein, two recently reappraised
47 species from the Jurassic, and five additional species from the Cretaceous and Miocene.

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48

49 Results. Contrary to prior assumptions, our results indicate that none of the Jurassic taxa
50 investigated belong to Asteroiidae or any other modern families, and instead represent stem-
51 forcipulatids. Furthermore, our phylogenetic results suggest that Asteroiidae likely originated
52 during the late Cretaceous. Our findings highlight a greater early diversity within the
53 Forcipulatacea than previously presumed, challenging existing perceptions of the evolutionary
54 history of this significant clade of marine invertebrates.

55

56 Introduction

57

58 The Forcipulatacea is one of the major clades within the Neoasteroidea, with about 400
59 described extant species (Mah and Blake, 2012). The crown-group Neoasteroidea is believed to
60 have diversified after the Permo-Triassic mass extinction (Blake, 1987; Gale, 1987). However,
61 the earliest unambiguous forcipulatacean fossils are dated from the Lower Jurassic (Blake,
62 1990), resulting in a major gap in the understanding of the early history of the group. The
63 Triassic group Trichasterospida has been interpreted as stem Forcipulatacea (Blake 1987; Blake
64 & Hagdorn 2003, Blake et al. 2006), but recent discoveries have challenged the phylogenetic
65 status of trichasteropsids (Thuy et al., 2017; Villier et al., 2018). The early history of the
66 Forcipulatacea and its relationship with Triassic groups is important for understanding the
67 history of the Neoasteroidea, because of the ongoing uncertainty concerning their phylogenetic
68 position within the Neoasteroidea and relationships with Palaeozoic taxa (e.g., Blake, 1987;

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trichasteropsids throughout

69 Gale, 1987; Blake & Hagdorn, 2003; Gale, 2011a; Mah & Foltz, 2011a; Mah & Foltz, 2011b;
70 Mah & Blake, 2012; Blake & Mah, 2014; Linchangco et al., 2017). The interpretation of fossils
71 and the timing of character appearances are central to the debate of clade definition and deep
72 phylogenetic relationships.

73 The superorder Forcipulatacea comprises three orders: the Forcipulatida, (about 250
74 extant species); the Brisingida, (about 110 extant species) and the Zoroacallida (less than 40
75 extant species) (Mah & Blake, 2012). The systematics of the Forcipulatacea has changed
76 multiple times over the course of the last century, with some authors recognizing many families
77 (e.g. Asteriidae, Zoroasteridae, Pedicellasteridae, Pycnopodiidae, Labidiasteridae,
78 Neomorphasteridae; e.g. Clark & Downey, 1992; Clark & Mah, 2001) while others accepted
79 only two (Asteriidae, Zoroasteridae; e.g. Fisher, 1928).

80 The majority of extinct forcipulatid species have been placed in modern families. Most
81 of these have been assigned within the Asteriidae (e.g., *Hystrixasterias hettangiurnus* Blake,
82 1990; *Germanasterias amplipapularia* Blake, 1990; *Polarasterias* Rousseau & Gale, 2018 (in
83 Rousseau et al., 2018); *Savignaster villieri* Gale, 2011b). including the monospecific subfamily
84 Dermasterinae (containing the fossil genus *Dermaster*) (Hess 1972). Outside of the Asteriidae,
85 only three extinct species have been interpreted as Stichasteridae (*Argoviaster occultus* Hess,
86 1972; *Pegaster stichos* Blake & Peterson, 1993; and *Atalopegaster gundersoni* Blake &
87 Guensburg, 2016) and one interpreted as a Pedicellasteridae (*Afraster scalariformis* Blake et al.,
88 1996). There are no extinct taxa known for the Heliasteridae, but exceptionally preserved
89 specimens attributed to the extant species *Heliaster microbrachius* have been reported from the
90 Pliocene of Florida, USA (Jones & Portell, 1988). More generally, Cenozoic fossils are rare and
91 have all been compared to extant genera (Barker & Zullo, 1980; Blake & Zinsmeister, 1988;

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92 Blake & Aronson, 1998; Palópolo et al., 2021). The Brisingida are very rare in the fossil record,
93 and are only known from the Miocene of Japan (Yamaoka, 1987).

94 Recent phylogenetic analyses allowed reappraisal of the historic classifications of the
95 Forcipulatacea (Mah, 2000; Foltz et al., 2007; Mah & Foltz, 2011a; Mah et al., 2015). The most
96 comprehensive molecular-based phylogenetic analysis of Mah & Foltz (2011a) identified 4 main
97 clades: Asteriidae, Stichasteridae, Zoroasteridae and Brisingida. They further suggested that the
98 family Pedicellasteridae is polyphyletic, that Labidiasteridae is a synonym of Heliasteridae, and
99 that Neomorphasteridae is a synonym of Stichasteridae (Mah & Foltz, 2011a). More recently, a
100 study based on morphological data by Fau & Villier (2020) found congruent evidence supporting
101 the same four major extant subclades within the Forcipulatacea as Mah and Foltz (2011a). In a
102 second study focused on extant Zoroasteridae and their fossil relatives, Fau & Villier (2023)
103 suggested the Mesozoic Terminasteridae are not monophyletic, instead arguing to resurrect and
104 expand the Order Zorocallida Downey, 1970 to include both the crown group Zoroasteridae, and
105 fossil stem-group taxa. These recent phylogenetic analyses have greatly restricted the definition
106 of the Asteriidae and suggest greater phylogenetic and taxonomic diversity within the
107 Forcipulatacea than currently expressed in the literature (Mah & Foltz, 2011a; Mah et al., 2015;
108 Fau & Villier, 2020).

109 To better understand the origin and early diversification of this major superorder, we
110 present a taxonomic re-evaluation of six well-preserved Jurassic forcipulatacean taxa, including
111 the earliest crown-group members. Further, we revise two species: *Forbesasterias gaveyi* gen.
112 nov. (Forbes, 1850) and *Marbleaster spiniger* gen. nov. (Wright, 1880) and redescribe
113 *Argoviaster occultus* Hess 1972 and *Dermaster boehmi* de Loriol, 1899. Using the largest
114 morphology-based character matrix ever constructed for the Forcipulatacea (Fau & Villier,

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“historic” classifications ? Please give references.

Commenté [ST7]: corresponding to the families and order...

Commenté [ST8]: of another family ? Please precise.

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Commenté [ST10]: And before all describe two new genus !

115 2020), we apply Bayesian tip-dating phylogenetic methods (Warnock and Wright, 2020; Wright
116 et al., 2021) to simultaneously co-estimate phylogenetic relationships and divergence times
117 among fossil and extant Forcipulatacea. In addition to the six species revised herein, we also
118 sampled seven other fossil forcipulatacean taxa (from Jurassic to Miocene) in our tip-dating
119 analysis to assist the estimation of divergence times. Finally, the results of our divergence dating
120 analysis allows us to investigate questions surrounding the origination and diversification of
121 major forcipulatacean subclades, including the species-rich Asteroiidae.

Commenté [ST11]: selected

Commenté [ST12]: species ?

123 Material and Methods

124 Taxon sampling

125
126 The analysis focuses on the most completely known extinct Jurassic taxa to ensure the
127 maximum number of characters could be scored. The six Jurassic taxa reappraised are
128 represented by 37 specimens (see material examined under systematic palaeontology). The
129 phylogenetic analysis is based on our reappraised descriptions and taxonomic revision. Seven
130 extinct species were added in the phylogenetic analysis in order to obtain a more comprehensive
131 temporal coverage of the Forcipulatacea. Two recently reappraised Jurassic taxa: *Psammaster*
132 *davidsoni* reappraised in Fau et al. (2020); *Terminaster cancriformis* reappraised in Fau & Villier
133 (2023); and three Cretaceous taxa: *Pegaster stichos* Blake & Peterson, 1993; *Cretasterias*
134 *reticulatus* Gale & Villier, 2013; *Viridisaster guerangeri* Fau & Villier 2023; and two Cenozoic
135 taxa: *Zoroaster marambioensis* Palópolo et al., 2021; *Brisingella* sp..

137 Terminology

138

139 The anatomical descriptions follow the terminology outlined by Fau and Villier (2018;
140 2020). Anatomical terms and abbreviations from Fau and Villier (2018; 2020) are in italics in the
141 text. We use conventional terms for the orientation of the specimen: abactinal (adoral) versus
142 actinal (oral); proximal versus distal; and abradial versus adradial. In the literature, the size of an
143 individual is commonly given with the two measures “r” and “R”, with r corresponding to the
144 distance between the centre of the disc and the edge of the disc, and R corresponding to the
145 distance between the centre of the disc and the tip of the arm.

146

147 **Nomenclatural act**

148

149 The electronic version of this article in Portable Document Format (PDF) will represent a
150 published work according to the International Commission on Zoological Nomenclature (ICZN),
151 and hence the new names contained in the electronic version are effectively published under that
152 Code from the electronic edition alone. This published work and the nomenclatural acts it
153 contains have been registered in ZooBank, the online registration system for the ICZN. The
154 ZooBank LSIDs (Life Science Identifiers) can be resolved and the associated information viewed
155 through any standard web browser by appending the LSID to the prefix <http://zoobank.org/>. The
156 LSID for this publication is: [urn:lsid:zoobank.org:pub:6A43BD80-6C00-42C6-AFD4-
157 7C4A944396FD]. The online version of this work is archived and available from the following
158 digital repositories: PeerJ, PubMed Central SCIE and CLOCKSS.

159

160 **Phylogenetic Analysis**

161

162 We expanded the morphological character matrix of Fau & Villier (2020, 2023) to
163 sample a total of 13 fossil species (6 taxonomically re-evaluated herein) and 29 extant species for
164 120 characters (character matrix available in Supplementary Materials). The extant *Plutonaster*
165 *bifrons* and *Dactylosaster cylindricus* were also sampled in our matrix as outgroup taxa. In this
166 study, the character/taxon matrix of Fau & Villier (2020, 2023) was revised as followed: the
167 character states of character 50 (number of primary spines per adambulacral; character number
168 49 in Fau & Villier, 2020) were modified, as many extinct taxa possess three or four
169 adambulacral spines (i.e., state 0: 1 to 2 adambulacral spines; state 1: 3 to 4 adambulacral spines;
170 state 2: 5 and more adambulacral spines), and two characters were added (character 119 and 120;
171 see Supplementary Materials File). The modified character by taxon matrix was coded using
172 MESQUITE software (Maddison & Maddison, 2023). Our matrix is scored up to 75.1%
173 complete (24.9% of missing or non-applicable characters), among which the 13 extinct taxa are
174 scored up to 50% complete.

175 To simultaneously co-estimate phylogenetic relationships and estimate divergences
176 times, we conducted a tip-dated Bayesian phylogenetic analysis incorporating the fossilized
177 birth-death process (FBD) (Stadler, 2010; Heath et al., 2014; Gavryushkina et al., 2014; Wright,
178 2017a; Warnock and Wright, 2020; Wright et al., 2021). Bayesian phylogenetic methods using
179 FBD models leverage both morphological and stratigraphic age information from the fossil
180 record (Barido-Sottani et al., 2020; Wright et al., 2021), which can then be combined with data
181 from extant taxa to generate time-calibrated phylogenies containing both fossil and extant
182 species. Tip-dating approaches using the FBD process provide a more coherent framework for
183 dating lineage divergences than node-based approaches (Heath et al., 2014), and have also been

184 shown to improve phylogenetic inferences involving fossil taxa compared to undated approaches
185 (Barido-Sottani et al., 2020; Mongiardino Koch et al., 2021). Moreover, Bayesian tip-dating
186 approaches can be used to directly test macroevolutionary hypotheses about character evolution
187 (Wright, 2017b; Wright et al., 2021), patterns of clade diversification (Paterson et al., 2019), and
188 investigate the evolutionary origin of major clades (Wright and Toom, 2017; Thuy et al., 2022).

189 We applied the sampled-ancestor implementation of the FBD model (Gavryushkina et al.,
190 2014), and placed broad, uniform priors on FBD parameters for diversification, extinction, and
191 fossil sampling. Fossil ages were assigned uniform distributions based on their occurrences in
192 geologic stages. An abundance of paleontological evidence points to a post-Permian origin of the
193 Neoasteroidea (e.g., Blake, 1987; Gale, 1987; Villier et al., 2018). To incorporate this
194 information while allowing for possible alternatives, we placed a prior distribution on the tree
195 age that spans the Permo-Triassic boundary (~240 Ma—260 Ma). Morphological character
196 evolution was modelled using a variant of the simple Mk model (Lewis, 2001) that accounts for
197 ascertainment bias and allows for morphological rates to vary among characters according to a
198 lognormal distribution (Wagner, 2012; Wright, 2017). To account for rate variation among
199 lineages throughout the tree, we applied an uncorrelated morphological clock where branch rates
200 vary according to an independent gamma rates model (Lepage et al., 2007).

201 To assist the analysis, we applied a topological constraint to the taxon *Brisingella* sp., a
202 taxon so poorly preserved that only 9 characters could be scored. *Brisingella* sp. possesses clear
203 characters showing its affinities with the Brisingida, such as a large number of arms (9 arms), the
204 presence of a rigid circular oral ring and the shape of its ambulacral (robust and hourglass). We
205 followed the work of Zhang et al. (2024), which considers *Brinsigella* sp. as the sister taxa to the

206 Freyellidae. Although poorly preserved, *Brinsigella* sp. is the only fossil of Brinsigida known to
207 date and was included here to sample fossil representatives for all forcipulataceans orders.

208 Bayesian inference of phylogeny was estimated using Markov chain Monte Carlo
209 (MCMC) simulation in MrBayes 3.2.6 (Ronquist et al., 2012). Two MCMC runs with four
210 chains were run for 80,000,000 generations. Chains were sampled every 500 generations and the
211 first 25% sampled were discarded as burn-in. Chains reached an average deviation split
212 frequency of less than 0.01. Convergence was assessed using Tracer 1.7.1 (Rambaud et al.,
213 2018), parameters attained effective sample sizes (ESS) >100, with all but one >1,000, and
214 potential scale reduction factors (PSRF) of ~1.0. The character/taxon matrix and MrBayes script
215 are available in the Supplementary Materials.

216

217 **Institutional Abbreviations:**

218

219 NHMUK, Natural History Museum, London, United Kingdom, formerly the British Museum of
220 Natural History (BMNH).

221 NMB, Naturhistorisches Museum Basel, Basel, Switzerland.

222 SMNS, Staatliches Museum für Naturkunde Stuttgart, Stuttgart, Germany.

223

224 **Systematic paleontology**

225

226 ASTEROIDEA de Blainville, 1830

227 FORCIPULATACEA Blake, 1987

228 *Forbesasterias* gen. nov.

229

230 **Type species.** *Uraster gaveyi* Forbes, 1850.

231 **Derivation of name.** In honor of Edward Forbes, who described this specimen in 1850.

232 **Diagnosis.** As for species, by monotypy.

233

234 *Forbesasterias gaveyi* (Forbes, 1850)

235 Fig. 1A, 2-3

236

237 1850 *Uraster gaveyi* Forbes decade III pl. II

238 1854 *Uraster gaveyi*, Forbes p. 90

239 1863 *Uraster gaveyi*, Wright, p. 100–102, pl. 1, fig. 1a, b

240 1870 *Uraster gaveyi*, Wright, p. 163

241 1876 *Uraster gaveyi*, Quenstedt, p. 85, pl. 93, fig. 29

242 ? 1935 *Asterias gaveyi*, Mercier, p. 47, pl. 2, fig. 18a-b

243 1966 *Asterias? gaveyi*, Spencer & Wright p. U75, fig. 66.1

244 1972 *Asterias? gaveyi*, Hess p. 32

245 1993 “*Asterias*” *gaveyi*, Lewis, p. 48

246 1996 “*Asterias*” *gaveyi*, Blake, p. 179

247 2011 “*Asterias*” *gaveyi*, Gale p. 58, text-fig. 24A-C

248

249 **Type specimen.** NHMUK PI E 1638, holotype.

250

251 **Type locality.** Mickleton tunnel, near Chipping Campden, Gloucestershire, England;

252 *Capricornus* zone, Pliensbachian, Early Jurassic (Forbes, 1850; Wright, 1863).

253

254 **Material examined.** NHMUK PI E 1638, holotype.

255

256 **Diagnosis.** Forcipulatacean sea star with five arms, compressed ambulacrals and adambulacrals,

257 and quadriserial ambulacral pores. Adoral carina composed of four elongated adambulacrals,

258 bearing one or two spines each. Other adambulacrals bearing four spines. Actinal plates present.

259 Straight pedicellariae present, regular straight pedicellaria approximately 1 or 2 mm long, and

260 giant straight pedicellaria approximately 5 mm long. Giant straight pedicellariae located around the

261 mouth in interbrachial areas. Abactinal spines short and numerous with ornamented tips.

262

263 **Description.** The holotype exposes the actinal face of a large individual, $R > 99$ mm, $r = 26$ mm,

264 with four arms preserved (Fig. 1). The specimen was slightly flattened during preservation

265 processes. The ambulacral grooves are probably wider than they would have been in life. Some

266 portions of the arms are more damaged than others, with disarticulated ambulacral heads in some

267 part of the arms. Ambulacral grooves are less flattened and therefore narrower around the mouth

268 frame. The disc and the structure of the wall skeleton are not exposed.

269 Ambulacrals are compressed lengthwise. The head is slightly broader than the shaft and

270 the teeth are present along the entire width of the ambulacral head (Fig. 2A-B, 3A-B). The *actam*

271 and the *furrow* are well defined (Fig. 3B), no *wings* on the ambulacral bases. Four tube feet rows

272 per ambulacral groove. The distalmost part of one of the arms is crushed, allowing observation

273 of the proximal side of a few ambulacrals and adambulacrals. The most distal ambulacrals are

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274 rather straight, their actinal edge being straight to slightly concave, as in many Forcipulatacea.

275 Arching of the actinal edge of proximal ambulacrals is not visible due to preservation.

276 Adambulacrals are compressed lengthwise, bearing 4 spines each, arranged in a
277 transverse row. The adambulacral spines are short and thick, slightly flattened at the extremity,
278 but longer than the actinal spines. The spines do not seem to have glassy trabeculae or any
279 ornamentation, instead are composed of undifferentiated labyrinthic stereom.

280 The adoral carina is composed of at least 4 elongate adambulacrals per ambulacral side
281 (Fig. 2). Adambulacrals of the adoral carina have a triangular shape, whereas they are more or
282 less square in others. They are also, at least, 1.5 times longer than the other adambulacrals with
283 the two most proximal adambulacrals being the longest. There are only one or two spines on the
284 adambulacrals of the adoral carina. The oral ossicles are short and bear at least two spines each.

285 Actinal and possibly marginal plates are visible, most of them still bearing spines.

286 Actinals overlapping each other. Actinal rows can be distinguished from marginal rows by
287 ending before the terminal ossicle. There are at least two rows of actinals, very likely more, but it
288 is impossible to count actinal rows near the interradius where the number is usually at its
289 maximum. The longest actinal row ends at least six millimetres before the end of the longest
290 preserved arm tip, even though the actual tip is missing. This means that the actinals were an
291 important part of the body wall. Actinal plates are small and stout. They are arranged to form
292 regular lateral and longitudinal rows. Some actinals have a small central *psas* (i.e. primary spine
293 attachment structure, also called “*pustule*” in Fau & Villier, 2020) supporting a primary spine,
294 but this structure is not visible on every plate.

295 A row of bigger, slightly triangular plates, is visible in the interrarial area. These plates
296 are morphologically differentiated compared to the actinal plates and are here assumed to be

297 marginal plates. Morphological differences between actinal and marginal plates reduces distally
298 along the arm. Due to preservation, it is impossible to determine if a second marginal series is
299 present. Abactinals and carinals cannot be observed. Short and slender spines, with ornamented
300 extremities, may represent spines of dorsal plate series.

301 Giant straight pedicellariae (up to 5 mm long) are found proximally in the interradial
302 space around the mouth and inside the ambulacral groove (Fig. 2C, 3C-E). The giant straight
303 pedicellariae are made of two slender and long valves that broaden at the base. Smaller straight
304 pedicellariae, with two slender regular valves that are slightly flattened at the extremity, occur
305 more distally along the arms. No crossed pedicellariae are recognized.

306
307 **Remarks.** The identification of the specimen NHMUK PI E 3339 as “*Uraster*” *gaveyi* is
308 uncertain because numerous morphological differences are apparent with the holotype. The
309 description is therefore based on the holotype only.

310 Mercier (1935) attributed disarticulated body wall skeleton ossicles materials from the
311 Sinemurian of Normandy, France to “*Asterias*” *gaveyi*. However, this material cannot be
312 reliably compared with the holotype and is therefore excluded.

313 The distinctive giant pedicellariae of *Forbesasterias gaveyi* gen. nov. are noticeable, but
314 giant pedicellariae also occur in modern forcipulataceans taxa. For instance, the asteriid
315 *Notasterias armata* possesses giant crossed pedicellariae on its abactinal surface. Large straight
316 pedicellariae are also found in Zoroasteridae.

317 *Forbesasterias gaveyi* gen. nov. is clearly distinguished from other extant species of the
318 genus *Asterias* by the lack of crossed pedicellaria, that it shows no evidence of intermarginals,
319 and possesses four spines per adambulacral. Species of extant *Asterias* possesses both straight

a supprimé: A

321 and crossed pedicellariae, clearly distinguishable intermarginals and adambulacrals with one to
322 three spines.

323

324 *Marbleaster* gen. nov.

325

326 **Type species.** *Marbleaster spiniger* (Wright, 1880)

327 **Derivation of name.** For the Forest Marble Formation.

328 **Diagnosis.** As for species, by monotypy.

329

330 *Marbleaster spiniger* (Wright, 1880)

331 Fig. 1B, 4

332

333 1880 *Uraster spiniger* Wright p. 166-167, fig. 1; pl. XXI, fig. 1

334 1966 *Compsaster spiniger*, Spencer & Wright p. U74, fig. 65, 1c

335 1993 *Compsaster spiniger*, Lewis, p. 60

336

337 **Type specimen.** NHMUK PI E 1642, holotype.

338

339 **Type locality.** near Rode (formerly Road), Somerset, United Kingdom; Forest Marble

340 Formation, Bathonian, Middle Jurassic (Wright, 1880). Wright (1880) refers to the locality as

341 “near Road, Wilts”, however the village of Rode sits on the Wiltshire to Somerset boarder and is

342 now regarded as part of the latter county.

343

344 **Material examined.** NHMUK PI E 1642, holotype.

345

346 **Diagnosis.** Forcipulatacean sea star with five short arms. At least three spines per oral ossicle,
347 short adoral carina composed of one or two adambulacrals only. Ambulacrals and adambulacrals
348 compressed, 3 to 4 spines per adambulacral. Ambulacrals with an extended *crest* on the
349 ambulacral's *head*, and a well-defined *furrow* on the ambulacral's *shaft*. Body wall plates present
350 with keyhole-shaped primary spine attachment structure (*psas*) in the interradial area of the disc.
351 Body wall ossicles bearing many long and slender spines, made of glassy trabeculae. Straight
352 and crossed forcipulate pedicellariae of the same size sparsely distributed across the body.

353

354 **Description.** The specimen has a diameter of about 4 cm. Only the actinal surface is visible. The
355 body is flattened, two arms are broken, and most of the ambulacral grooves are covered by
356 sediments. The body wall skeleton is not accessible. Some plates in the interradial area can be
357 interpreted as actinal plates.

358 The oral frame is characterised by five pairs of long and narrow oral ossicles, each
359 bearing at least one spine directed proximally and one or two spines oriented actinally. Short
360 adoral carina composed of the first proximal adambulacral only, with some second
361 adambulacrals in contact but not compressed in width. Adambulacrals of the adoral carina are
362 narrower and longer than other adambulacrals. They possess only one *psas*, instead of 3 as the
363 rest of the adambulacrals along the arms.

364 Ambulacral grooves are partially covered by spines and sediments. Ambulacrals are
365 compressed in length, the *head* larger than the *shaft*, slightly hourglass-shaped. The *actam* and
366 the *furrow* (on the ambulacral *shaft*) are well defined. No *wings* on the ambulacral *bases*. A

367 *crest*, similar to the ambulacral crest of extant asteriids (Fig. 4; Fau & Villier, 2020) is present on
368 the head and it is tilted in a proximal direction.

369 The adambulacrals bear 3 to 4 relatively thick and long spines (2 mm) with ornamented
370 tips and attached to a *psas*. The adambulacrals are irregular in size, highly compressed in length,
371 and wider than high.

372 The body wall plates are present in the interradial area. At least 2 rows of body wall
373 plates can be identified. They are likely to be actinal plates because of their small size and
374 placement. One of these plates possesses a keyhole-shaped *psa*. The arm plates are covered by
375 long and conic spines made of glassy trabeculae.

376 Both straight, duck-billed pedicellariae and crossed forcipulate pedicellariae are present
377 (Fig. 4C). They are randomly distributed between the spines across the surface of the specimen.
378 Straight and crossed pedicellariae are of similar size (1 mm). The crossed pedicellariae are
379 relatively large, some are as long as half the length of the spines. Crossed pedicellariae are
380 similar in shape to those of modern Stichasteridae and Asteridae (e.g. Fisher 1928, 1930, Clark
381 & Downey, 1992), and are randomly distributed on the abactinal surface. There is no evidence of
382 wreath organs.

383

384 **Remarks.** “*Uraster*” *spiniger* was originally classified with in the asteriids by Wright (1880).
385 However, he expressed doubt about the systematic position of his new species: “This starfish
386 differs so much from the other fossil species of the genus *Uraster* that it may possibly prove to
387 be the type of a new genus, when more details are learned anent the anatomy of the skeleton by
388 the discovery of new materials.” (Wright, 1880, p. 167). Spencer & Wright (1966) assigned
389 “*Uraster*” *spiniger* to the genus *Compsaster*. However, the type species of the genus *Compsaster*

390 *formosus* Worthen & Miller, 1883 from the Carboniferous of Illinois differs in many aspects
391 from “*Uraster*” *spiniger* Wright, 1880. These differences have been outlined by Blake (2002, p.
392 363): “Although the type specimen of the Jurassic species *Compsaster spiniger* is imperfectly
393 preserved, it appears readily assigned to the surviving Asteriidae, a family known from the
394 beginning of the Jurassic (Blake, 1990), well before the Bathonian occurrence of *C. spiniger*.
395 The Compsasteridae therefore here is restricted to the type species.” As outline by Wright
396 (1880), *Marbleaster spiniger* gen. nov. exhibits distinctive characters that warrant the
397 establishment of a novel genus to house this species. These include the presence keyhole-shaped
398 *psas*, and the presence of straight and crossed pedicellariae of equal size uniformly distributed
399 across its actinal surface.

400

401 *Dermaster boehmi* de Loriol, 1899

402 Fig. 5

403

404 1899 *Dermaster boehmi* de Loriol, p.1-6, pl. 1, fig. 1

405 1972 *Dermaster boehmi*, Hess p. 32-36, text-fig. 3, 31-39, 89; pl. 2, fig. 1, 3; pl. 3, fig. 1; pl. 4,
406 fig. 1; pl. 12, fig. 2

407 1973 *Dermaster boehmi*, Hess, p. 627

408 2011 *Dermaster boehmi*, Gale p. 60

409

410 **Type specimen:** Specimen illustrated by de Loriol, 1899, pl. 1, fig. 1 (assumed lost).

411

Commenté [ST14]: Can a neotype be defined ?

412 **Type locality.** Vögisheim, Mülheim, Baden-Württemberg, Germany; ferruginous layers in
413 limestone, Bathonian, Middle Jurassic (de Loriol, 1899). From the information provided by de
414 Loriol (1899; locality, age and geology), the holotype is assumed to come from the ferruginous
415 oolitic layer of the Hauptrogenstein-Formation (Bloos et al. 2006).

416

417 **Material examined.** Six well-preserved specimens described by Hess (1972): NMB M8985,
418 M10678, M9365, M9600, M10705, and M9168. All specimens originate from the village of
419 Schinznach, Canton of Aargau, Switzerland and were collected from the Upper Hauptrogenstein-
420 Formation (Upper Bajocian; Middle Jurassic).

421

422 **Diagnosis** (emended from Hess, 1972). Aboral skeleton reticulate, pore fields present but small.
423 Body wall ossicles cruciform to triangular, covered by small granule-like spines. One or two
424 *psas* on primary interradials and primary radials. Primary radials overlapping the primary
425 interradials. Adoral carina present, composed of the two most proximal adambulacrals.
426 Ambulacrals compressed, ambulacral pores biserial. Adambulacrals with 3 to 4 spines each.
427 Straight and crossed pedicellariae present. Crossed pedicellariae differentiated into two
428 morphotypes.

429

430 **Description.** The disc is composed of five primary radials and five primary interradials that are
431 arranged around a circlet of small abactinals and the primary central plate (Fig. 5A-B). The
432 primary central plate is approximal twice the size of the disc abactinals. The superomarginal
433 plate rows extend into the disc to join and to partially overlap the primary interradials. The most
434 proximal superomarginals of adjacent arms are in contact inter-radially on the disc. In NMB

435 M10678, only one of the enlarged primary radials overlaps the primary interradials directly. The
436 other four primary radials are slightly set distally and not in contact with the other primary
437 interradials. All disc plates are covered by little bumps, that indicate the former presence of
438 spines. NMB M10705 is an arm fragment, with what is likely to be a few plates from the disc.
439 One primary radial is visible and bears a relatively big *psas* in its center. The radial of NMB
440 M10705 is twice the size of those of NMB M10678 and clearly show two types of attachment for
441 the spines, a big central *psas* and some smaller bumps around the *psas*, as already described by
442 Hess (1972). At least two types of spines (i.e., primary spines attached on *psas*, and secondary
443 spines attached on *bumps*) are present. The madreporite is not preserved in any of the specimens
444 studied. In NMB M10678, a cavity remains in one of the primary interradials, where the
445 madreporite inserted (Fig. 4.5A-B). De Loriol (1899) and Hess (1972) described the madreporite
446 as a small swollen plate, that “lies near the edge of the interbrachial angle” (translation of de
447 Loriol, 1899 p. 5). Current observations and the description of de Loriol (1899) and Hess (1972)
448 agree with the conclusion that the madreporite was not fused with an interradial. A strong
449 relationship between one of the interradials and the madreporite is possible, as in the
450 Zoroasteridae in which the madreporite sets in a special cavity on the distal edge of an
451 interradial.

452 The arms are composed of at least nine plates rows in addition to the adambulacral and
453 ambulacral columns: one row of carinals, one row of abactinals, two rows of marginals, and one
454 or two rows of actinals, on each side. Proximally, carinals, abactinals and superomarginals are
455 cruciform, but the articular processes get shorter distally (Fig. 5B, D). All plate surfaces are
456 granulated. The carinal row is regular along the arms. The carinals were formerly adorned by
457 many small spines, and at least one big stout rounded spine. Spines are better preserved in NMB

458 M10705: there are big and stout, blunt primary spines, and small acicular secondary spines.
459 There is no doubt that there were many secondary spines per plates because of their granulated
460 surfaces. Because of the number of primary spines preserved, it is also likely that each plate was
461 bearing one or more primary spines. However, *psas* are not observed on every plate.

462 Abactinals are small and irregular in shape. They are overlapped by both the carinals and
463 the superomarginals. Superomarginals are cruciform proximally, but of rather heterogenous
464 shape distally. Inferomarginals are smaller than the superomarginals. All bore many spines, at
465 least one primary spine and many secondary spines. The terminal ossicle is round and relatively
466 big, with a granular surface, probably bearing many spines too. At least one row of actinals is
467 present on the specimens studied. The actinals are small, with a cruciform to triangular shape, the
468 abactinal lobe tends to be reduced.

469 Ambulacrals are compressed, but not as much as in modern Asteriidae. Tube feet are
470 arranged in two rows in the ambulacral groove. The *head* of the ambulacrals are slightly longer
471 than the *shaft*, but symmetrical, the *furrow* on the *shaft* is well marked, and no wings on the
472 ambulacral *bases*. Adambulacrals are also compressed and bear a transverse row of at least 3,
473 maybe 4 spines per adambulacral. Adambulacral spines are conical, long and slender. They are
474 the longest spines present. These long and slender spines are preserved only around the
475 adambulacrals. There are no signs of long primary spines on the actinals or on the
476 inferomarginals. The adoral carina is short and composed of only the two most proximal
477 adambulacrals.

478 Both straight and cross pedicellariae are preserved in NMB M10705 and M10678.
479 Straight pedicellariae are similar to straight pedicellariae of extant taxa and are the most visible
480 on the actinal surface of NMB M10678-B. It is possible to recognize two types of crossed

481 pedicellariae. In NMB M10678, small crossed pedicellariae are abundant, especially between
482 around the marginals and the actinals. In NMB M10705, on the other hand, crossed pedicellariae
483 are larger and more robust.

484

485 **Remarks.** *Dermaster boehmi* seems to present two different types of crossed pedicellariae. Only
486 the “robust” crossed pedicellariae were described and illustrated by Hess (1972, Fig. 33, 35-37).
487 *D. boehmi* is not the only species to present different morphotypes of pedicellariae. For instance,
488 *Pisaster ochraceous* has two types of straight pedicellariae while *Pedicellaster hypernotius* has
489 two types of crossed pedicellariae (Fau & Villier, 2020). Even if this is rare among the
490 Forcipulatacea, this is not a unique case of multi-pedicellariae morphotypes. However, until this
491 polymorphism is found in other fossil taxa, this should be regarded as an autapomorphy of *D.*
492 *boehmi*.

493

494 *Argoviaster occultus* Hess, 1972

495 Fig. 6

496

497 1972 *Argoviaster occultus* Hess p. 27-32, text-fig. 29-30; pl. 9, fig. 2; pl. 10-11; pl. 12, fig.1; pl.
498 13, fig.1; pl. 14, fig.1

499

500 **Type specimen.** NMB M8977, holotype.

501

502 **Type locality.** Schinznach, Canton of Aargau, Switzerland; Upper Hauptrogenstein-Formation,
503 Upper Bajocian, Middle Jurassic (Hess, 1972).

a supprimé: .

505

506 **Material examined.** NMB M8977, holotype; NMB M9359, NMB M9362, NMB M9366 NMB
507 M9344/1-2, NMB M9360/1-2, NMB M9361/1-2, NMB NMB M9364/1-2, NMB M10676,
508 paratypes; NMB M9460, NMB M9465, NMB M9469, NMB M9475, NMB M9479, NMB
509 M9480/1-2, NMB M9481, NMB M9482, NMB M9483, NMB M9487, NMB M9489, NMB
510 M9505, NMB M9506/1-2, NMB M9514/1-2.

511

512 **Diagnosis** (emended from Hess, 1972). Forcipulatacean sea star with five arms. Ambulacrals and
513 adambulacrals compressed. Ambulacral pores quadriserial. Adambulacral with three short spines
514 each. Plates of the arms arranged in longitudinal and transverse rows, with a small papular field
515 at each corner. Body wall ossicles triangular to cruciform, with reduced ornamentation of
516 granules, sometimes with one central *psas*, and occasionally a central cavity. Straight and
517 crossed forcipulate pedicellariae present.

518

519 **Description.** The holotype, NMB M8977 is a distal part of an arm, showing mostly the
520 ambulacral groove. NMB M9366 is composed of four arms partially preserved in abactinal view,
521 and a few ossicles of the fifth arm still embedded in the matrix. NMB M9362 has five partially
522 preserved arms, the body wall skeleton is mostly missing, so that the oral frame and the
523 ambulacral skeleton can be observed from the inside.

524 The structure of the wall skeleton is visible on one arm of NMB M9366, but because the
525 specimen is flattened, the arms appear larger than they would have been in life. All the skeleton
526 arm plates in NMB M9366 are small and triangular, the surface is finely granulated with no *psas*.
527 The central plates row is assumed to be the carinal row (Fig. 6A-B), because of its central

528 position, and because the carinals overlap their abactinal neighbours on each side of the arm. On
529 each side of the carinals, there is at least one row of abactinals, with small plates intercalated in
530 between the abactinal and carinal rows. All arm plates series are similar in shape and size and it
531 is difficult to differentiate abactinal plates from the marginal plates. There are possibly one row
532 of superomarginals and one row of inferomarginals on each side of the arm. The second-best
533 preserved arm of NMB M9366 shows part of the carinal row proximally, overlapping some
534 abactinals. The rest of the carinals and abactinals have been removed, exposing the ambulacrals,
535 adambulacrals and some actinals and maybe inferomarginals that are cut transversally (Fig. 6A-
536 B). There are at least 3 or 4 rows of actinals.

537 Wall skeleton plates in NMB M9362 look different in shape compared to the triangular
538 plates observed in NMB M9366, but the differences could be explained by the different views
539 offered by the two specimens. Ambulacrals and adambulacrals are easily recognizable in NMB
540 M9362 (Fig. 6C-D), but the wall skeleton plates are more difficult to recognise because they
541 were scattered by taphonomic disarticulation. Hess (1972) recognized difficulties in the
542 identification of superomarginals in M9362. Assuming that the wall skeleton plates of *A.*
543 *occultus* follow the Forcipulatid Plating Rules (Gale, 2011a), the homologies of the actinal,
544 inferomarginal, superomarginal, abactinal and carinal series can be recognized from their relative
545 position, shape and number of articulation areas. As the body wall plates overlap one another,
546 they present a number of articulation areas on their external and internal faces. When looking at
547 the internal faces, carinals and superomarginals should exhibit three articulation areas, the
548 inferomarginals only two, and the actinals and abactinals one or two. Carinals are cruciform,
549 abactinals seems to be triangular in radial cross-section, or rod-like in actinal view.
550 Superomarginals are also cruciform, but higher than long, with a well-developed actinal lobe.

551 Inferomarginals on the contrary have a more developed abactinal lobe. Actinals seems to be
552 more or less cruciform. Two rows of actinals are visible in the holotype. They strongly overlap
553 each other and bear one *psas* per plate (Fig. 6E). All three specimens have wall skeleton plates
554 with a central cavity, which is a unique feature in the Forcipulatacea.

555 Ambulacrals are highly compressed and gently curved to accommodate four alternate
556 tube feet rows in the ambulacral groove. The ambulacral crest is tilted proximally. There is no
557 wing on the ambulacral bases. A furrow is present and well-marked on the ambulacral shaft. The
558 adambulacrals are compressed as well and bear 3 spines each.

559 The specimen M9362 partially exhibits the oral frame ossicles. Only the *ramus* of the
560 orals is visible. Several spines are preserved in the center of the circle formed by the oral frame,
561 but it is not possible to count or estimate the number of spines per orals. First ambulacrals have a
562 shape similar to modern Asteriidae with a long head, and a long, well-developed, but not high,
563 proximal process. The odontophores are square and were probably connected to both the orals
564 and first ambulacrals with the articulation areas *poda* and *doda* clearly separated. The crater
565 seems to be present.

566 Pedicellariae were not found in NMB M9362. In NMB M9366, Hess (1972) described
567 remains of pedicellariae, scattered between the body wall ossicles. In the holotype, both straight
568 and crossed forcipulate pedicellariae are present along the ambulacral groove (Fig. 4.7).

569

570 **Remarks.** There is no mention of crossed pedicellaria in the original description of the species
571 by Hess (1972), but they are present at least in the holotype. In addition, the present description
572 provides more detail about the structure of the body wall skeleton (Fig. 6).

573

574 *Germanasterias amplipapularia* Blake, 1990

575 Fig. 7

576 1990 *Germanasterias amplipapularia* Blake p. 103-123, fig. 1-2

577 2011 *Germanasterias amplipapularia* Gale p. 57

578

579 **Type specimen.** SMNS 18869a-b.

580

581 **Type locality.** Göppingen, Baden-Württemberg, Germany; *Schlotheimia angulata* Zone, late
582 Hettangian, Early Jurassic (Blake, 1990).

583

584 **Material examined.** SMNS 18869a-b, holotype.

585

586 **Diagnosis** (emended from Blake, 1990). Forcipulatacean with body wall skeleton consisting of a
587 carinal series, one row of abactinals (and associated smaller plates) on each side of the carinal
588 series, two marginal and three actinal series. Carinals and marginals alternating between spine
589 bearing and non-spine bearing plates. Only one large primary spine per carinal or marginal with
590 spines. Abactinals digitate, arranged in regular transverse and longitudinal rows and separated by
591 relatively large papular area. Ambulacrals compressed, podial pores quadriserial. Adoral carina
592 composed of the first 4 to 5 adambulacrals. Adambulacrals compressed, bearing 4 spines each.
593 Straight duck billed pedicellariae present on abactinal surface only, very abundant.

594

595 **Remarks.** For complete description, see Blake (1990). Blake (1990) described the adambulacrals
596 of *Germanasterias amplipapularia* as weakly carinate proximally. In the literature,

597 adambulacrals of forcipulatacean sea stars are described as carinate if they possess an adradial
598 extension (Fau & Villier, 2018). Alternate carinate and non carinate adambulacrals are a
599 synapomorphy of the Zoroasteridae (Fau & Villier, 2020). Some other forcipulatacean taxa (e.g.,
600 *Heliaster*, *Asterias*) can have alternating sized adambulacrals, with generally the larger sized
601 adambulacrals bearing one more spine than the smaller sized adambulacrals. However different
602 sized adambulacrals cannot be considered homologous with the adradial extension of the
603 Zoroasteridae. The scoring of *G. amplipapularia* for phylogenetic study follows the scoring
604 system of Fau & Villier (2020), considering carinate adambulacrals absent on both *G.*
605 *amplipapularia* and *H. hettangiurnus*.
606
607 *Hystrixasterias hettangiurnus* Blake, 1990
608 Fig. 8
609
610 1990 *Hystrixasterias hettangiurnus* Blake p. 103-123, fig. 3-4
611 2011 *Hystrixasterias hettangiurnus* Gale p. 57-58, text-fig. 24D-G
612
613 **Type specimens.** NMB M9682, holotype; NMB M9681, NMB M9684-8, paratypes.
614
615 **Type locality.** Schechingen, Baden-Württemberg, Germany; Hettangian (Formation unknown),
616 Early Jurassic (Blake, 1990).
617
618 **Material examined.** NMB M9682, holotype; NMB M9684, NMB M9686

619 NMB M9687, NMB M9688, paratypes. Specimen NMB M9685 was missing from the NMB
620 collections, and assume to be lost.

621

622 **Diagnosis** (emended from Blake, 1990). Forcipulatacean sea star with body wall skeleton
623 consisting of carinals bordered on each side by up to three rows of abactinals, two rows of
624 marginals and two to three rows of actinals. Abactinals arranged in regular transverse and
625 longitudinal rows. Ambulacral moderately compressed, podial pores quadriserial. Adoral carina
626 composed of the first 3 adambulacrals. Adambulacrals with transverse series of four prominent
627 spines. Straight forcipulate pedicellariae present. Straight pedicellariae differentiated in two
628 types: clam-shaped pedicellariae present on abactinal surface only and duck billed pedicellariae
629 present on actinal surface.

630

631 **Remarks.** For complete description, see Blake (1990).

632 *Hystrixasterias hettangiurnus* possess two different morphotypes of straight pedicellariae,
633 relatively robust clam shell pedicellariae as illustrated by Gale (2011a, text-fig. 24D-G), and
634 modern-like duck billed pedicellariae (Fig. 8 C-D). *H. hettangiurnus* lacks keyhole-shaped *psas*
635 (synapomorphy of the Asteriidae) and possess a madreporite that is not fused with an interradial
636 (synapomorphy shared by the Asteriidae and the Stichasteridae).

637 Blake (1990) described the adambulacrals of *H. hettangiurnus* and *Germanasterias*
638 *amplipapularia* as weakly carinate respectively medially and proximally. Adambulacrals of *H.*
639 *hettangiurnus* are not here considerate carinate (see remarks for *Germanasterias*
640 *amplipapularia*).

641

Results

The main clades recovered in both the maximum credibility clade (MCC) tree (Fig. 9) and the 50% majority rule consensus (MRC) tree (Fig. 10) are: *Brisingida*, *Zorocallida*, *Stichasteridae* and *Asteriidae* (Fau & Villier, 2020; 2023). *Terminaster cancriformis* is found to be sister taxa to the *Zoroasteridae* in both the MCC tree and the majority rule consensus tree, retrieving the clade *Zorocallida* which is consistent with the results in Fau & Villier (2023). The position of *Hystrixasterias hettangiurnus*, *Germanasterias amplipapularia*, *Forbesasterias gaveyi* gen. nov., *Dermaster boehmi* and *Psammaster davidsoni* are found to be uncertain, but deeply rooted at the base of the Forcipulatida clade (Fig. 10). The position of *P. davidsoni* is compatible with the previous phylogenetic analysis sampling this taxon by Fau et al. 2020. *Dermaster boehmi* and *Psammaster davidsoni* are well supported as sister taxa (posterior probability = 0.75) (Fig. 10). Our results do not support *F. gaveyi*, *G. amplipapularia* and *H. hettangiurnus* as members of the family *Asteriidae*. *Argoviaster occultus* and *Marbleaster spiniger* are found to be the most derived of the Jurassic taxa reappraised here, as sister taxa to the Forcipulatida (Fig. 9).

The Forcipulatida is composed of the clade *Asteriidae* + *Stichasteridae*, the *Heliasteridae*, the extant species *Pedicellaster hypernotus* and the extinct taxa *Argoviaster occultus*, *Marbleaster spiniger* and *Pegaster stichos* (Fig. 10). Although the phylogenetic positions of these taxa are poorly supported (Fig. 9, 10), our results indicate none of the three extinct species belong to any extant families. Both *P. stichos* and *A. occultus* have been compared to the extant *Neomorphaster* and placed in the subfamily *Neomorphasterinae* (now synonymized with *Stichasteridae*; Hess, 1972; Blake & Peterson, 1993). However, our results do not support either of them as stichasterids. *Cretasterias reticulatus* Gale & Villier, 2013 is found to be the sister

Commenté [ST15]: + Heliasteridae and Zoroasteridae

Commenté [ST16]: Is it not the case of V guerangeri ?

Commenté [ST17]: unresolved

Commenté [ST18]: they rather seem to belong to the Forcipulatida

a supprimé: , 10

Commenté [ST19]: + the 5 other basal taxa

Commenté [ST20]: Is it worth keeping these nodes with so low support values ?

666 taxa to the clade formed by all extant asteriids. This position is compatible with either its
667 inclusion within the family Asteriidae or to be interpreted as a stem-Asteriidae. *Cretasterias*
668 *reticulatus* possesses 4 out of the 6 synapomorphies proposed by Fau & Villier (2020): char. 28
669 ambulacral with strongly arched abactinal profile; char. 29 the muscle insertion *lim* represent
670 more than 40% of the ossicle height and finishing under the *actam*; char. 68. Round *psas* absent
671 on the inferomarginals; char. 111. Wreath organ present. However, *C. reticulatus* does not have
672 any differentiated abactinals, which are the last two synapomorphies of the extant family (char.
673 80 abactinals differentiate with at least two level of plates; and char 82. abactinals, intercalary
674 inter-arc ossicles present).

675

676 Discussion

677

678 The Asteriidae Gray, 1840 was the first named family in the Forcipulatacea, initially
679 including all known forcipulataceans. The [number of](#) species assigned to this family have greatly
680 changed over time, as it was progressively split into several families. Fisher (1928, 1930) placed
681 many species into the family Asteriidae, including taxa that are currently classified in the
682 families Pedicellasteridae and Stichasteridae. Fisher's classification was followed by many
683 authors (e.g., Clark & Downey, 1992; Mah, 2000), until phylogenetic analysis, based on
684 molecular data, radically changed the circumscription of the Asteriidae (Mah & Foltz, 2011a).
685 Mah & Foltz's (2011a) phylogenetic hypothesis supports a restricted definition of the Asteriidae.
686 The Asteriidae and the Stichasteridae were retrieved as two distinct clades, and the family
687 Pedicellasteridae as polyphyletic (Mah & Foltz, 2011a). Even in its current and restricted
688 definition, the Asteriidae is still the most diverse family of all living forcipulatacean,

689 representing half of the species diversity of the group (Mah & Blake, 2012). The phylogenetic
690 hypothesis proposed by Fau & Villier (2020) based on morphological characters showed high
691 congruence with Mah & Foltz (2011a). Six morphological synapomorphies were found for the
692 clade Asteriidae (Fau & Villier, 2020): (i) a strongly arched abactinal profile of the ambulacrals,
693 (ii) a long muscle insertion *lim* that finishes under the *actam*, (iii) the absence of round *psas* on
694 the inferomarginals (spines attaching on keyhole-shaped *psas*, instead), (iv) abactinals
695 differentiate with at least two distinct plate shapes, (v) the presence of intercalary inter-arc
696 abactinals, and (vi) the presence of wreath organs. Only the presence of wreath organs is a non-
697 ambiguous synapomorphy. In the literature, the early Jurassic *F. gaveyi*, *M. spiniger*, *H.*
698 *hettangiurnus* and *G. amplipapularia* have all been considered, at some point, as members of the
699 family Asteriidae (Blake, 1990; Blake 2002). Our results no longer support them as members of
700 the clade Asteriidae (Fig. 9, 10).

701

702 Early Jurassic forcipulataceans

703

704 The three early Jurassic taxa *F. gaveyi*, *H. hettangiurnus* and *G. amplipapularia*, have
705 been interpreted as **asteriids**, mostly due to their highly compressed ambulacrals and
706 adambulacrals. Wright (1863- 1880 p. 101) wrote about *F. gaveyi*: “The structure of the
707 ambulacral skeleton, which is so admirably preserved in this fossil, removes all doubt as to its
708 true generic position and affinities”. Blake (1990) compared *H. hettangiurnus* and *G.*
709 *amplipapularia* to the Zoroasteridae and Asteriidae, stating that they are ‘intermediate in many
710 ways” (p. 104), but still concluded that they were both of asteriid affinities. At the time of
711 Blake’s (1990) publication, the family Asteriidae was not as restricted as it is today and

Commenté [ST21]: Asteriidae ?

712 comprised the subfamilies Stichasteridae and Pedicellasteridae. Therefore, Blake's (1990)
713 assumption of the phylogenetic position of *H. hettangiurnus* and *G. amplipapularia* is actually
714 congruent with our results.

715 The phylogenetic positions of *F. gaveyi*, *H. hettangiurnus* and *G. amplipapularia* in the
716 analysis suggest a new evolutionary history of the group. None of them possess any of the
717 synapomorphies of the Asteriidae as proposed by Fau & Villier (2020). Instead, they exhibit a
718 mix of plesiomorphic characters of the Forcipulatida and derived characters. Noticeable derived
719 characters shared by *F. gaveyi*, *H. hettangiurnus* and *G. amplipapularia* are the absence of wings
720 on the ambulacrals (character 35) and the high level of compression of the adambulacrals
721 (character 42). *F. gaveyi*, *H. hettangiurnus* and *G. amplipapularia* possess the following
722 plesiomorphic characters: (i) they have madreporites that are neither fused or imbricated with a
723 primary interradial (character 114, not applicable in *F. gaveyi*), and (ii) they do not have any
724 crossed pedicellariae. We cannot exclude the hypothesis that the absence of crossed pedicellariae
725 could be a taphonomic bias, as crossed pedicellariae are, in general, smaller than straight
726 pedicellariae. The absence of crossed pedicellariae could be a plesiomorphic characters within
727 the Forcipulatacea, as they are also absent in the Zorocallida, or it could be a convergent loss, as
728 crossed pedicellariae are present in other closely related Jurassic taxa, such as *P. davidsoni*, *D.*
729 *boehmi*, and *M. spiniger*.

730

731 ***Psammaster davidsoni* and *Dermaster boehmi***

732

733 The Middle Jurassic *D. boehmi* and the Late Jurassic *P. davidsoni* are found to be sister
734 taxa, and to be part of a polytomy in the 50% MRC tree (Fig. 10) along with the Early Jurassic

735 taxa. They share with the Early Jurassic taxa the following characters: numerous adambulacral
736 spines (four spines per adambulacral; character 50) and numerous actinal rows (at least three
737 actinal rows, character 56). Few extant forcipulatids possess more than 3 adambulacral spines (a
738 notable exception is the asteriid genus *Perissasterias*, which possesses up to 7 adambulacral
739 spines). It is thus noticeable that all the Jurassic forcipulatids analyzed here possess 3 to 4
740 adambulacral spines, grossly arranged in a transverse row, which could be the plesiomorphic
741 condition.

742 *P. davidsoni* and *D. boehmi* share with *G. amplipapularia* the following plesiomorphic
743 characters: presence of secondary spines on the abactinal skeleton (characters 62, 70 and 93), and
744 presence of only one row of abactinals between the superomarginals and the carinals (character
745 79). The presence of secondary spines is shared with Zorocallida. The presence of only one row
746 of abactinals is a plesiomorphic character, shared with *Labidiaster annulatus*, *Pedicellaster*
747 *hypernotius* and Zorocallida, and convergent in the stichasterid *Neomorphaster forcipatus*.

749 ***Argoviaster occultus* and *Marbleaster spiniger***

750
751 The Jurassic *A. occultus* and *M. spiniger* are found to be with the Cretaceous *Pegaster*
752 *stichos* higher in the tree, and belong without any doubt to the Forcipulatida. The close
753 relationship of *A. occultus*, *M. spiniger* and *P. stichos* to the Forcipulatida is supported by
754 characters 26 (average compression of the ambulacrals) and 28 (ambulacral arch slightly
755 concave). Average compression of the ambulacrals is shared by all forcipulatid and *G.*
756 *amplipaluarua*, *H. hettangiurnus*, *F. gaveyi*, *D. boehmi*, and *P. davidsoni*. No complete
757 ambulacrals were visible, thus character 28 was scored (1) slightly arch, or (2) strongly arch for

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760 both *A. occultus* or *M. spiniger*, pending check on availability of better-preserved fossils. Both *A.*
761 *occultus* and *M. spiniger* possess ambulacrals with small proximal tilting of the ambulacral's
762 head in the proximal directions (character 31). A small proximal tilting of the ambulacral's head
763 is a plesiomorphic character with most forcipulatid having rather pronounced proximal tilting.

764 Keyhole-shaped *psas* have been found on the actinals in *M. spiniger* (character 120) and
765 are also found on the actinals and inferomarginals of extant asteriids and stichasterids. Round
766 *psas* are also present on the inferomarginals of stichasterids but absent in asteriids. The presence
767 of a keyhole-shaped *psas* on the actinals of *M. spiniger* supports a derived position.

768 Of the Jurassic taxa, only *A. occultus*, *M. spiniger*, and *D. boehmi* possess both crossed
769 and straight pedicellariae. Only straight pedicellariae have been observe in *F. gaveyi*, *H.*
770 *hettangiurnus* and *G. amplipapularia*, and only crossed pedicellaria have been found in *P.*
771 *davidsoni*. However, the absence of straight pedicellariae in *P. davidsoni* is likely to be a
772 preservation bias (Fau et al., 2020). Except for *Cretasterias reticulatus* (Gale & Villier, 2013),
773 there is no evidence of the presence of wreath organs in any of the fossil taxa included here,
774 which is the only non-ambiguous synapomorphies of the Asteriidae. Wreath organs are a
775 concentration of crossed pedicellariae around some spines with dedicated muscles allowing the
776 wreath of pedicellariae to move up and down the spines (Lambert et al., 1984). Thus, none of the
777 Jurassic taxa can be assigned to crown Asteriidae.

778

779 **Other Jurassic forcipulataceans and their evolutionary significance**

780

781 Our phylogenetic analysis agrees with the suggestions of Gale (2011a) and Mah & Foltz
782 (2011a) that early Jurassic “asteriids” are not true Asteriidae. Our investigation suggests that

783 crown Asteroiidae were not yet present by the Early Jurassic. Historically, many fossil forms were
784 described as *Asterias* or synonyms of it during the late 19th and early 20th century, including
785 *Asteracanthion oolithicum* Terquem and Jourdy 1869 (Bathonian), *Asterias ranvillensis* Porte
786 1927 (Bathonian), and *Asterias delongchampsii* Morière 1878 (Oxfordian). Their assignment to
787 the genus *Asterias* appears unlikely, in view of the phylogeny of the Forcipulatacea, and they
788 need to be reappraised in the future. *Asterias? dubium* Whitfield 1877 (Jurassic) is a species
789 based on very poor material that do not allow for observation of morphological characters (Clark
790 & Twitchell, 1915), and is unlikely to represent the genus *Asterias* either (Clark & Twitchell,
791 1915; Whitfield, 1877; Whitfield, 1880). Two additional Jurassic fossil forms have been recently
792 interpreted as members of the Asteroiidae, *Savignasterias villieri* Gale, 2011b from the Oxfordian
793 of France and *Polarasterias janusensis* Rousseau & Gale, 2018 (in Rousseau, Gale & Thuy,
794 2018) from the Tithonian of central Spitsbergen. Although *Savignasterias villieri* and *P.*
795 *janusensis* were not included in this analysis, absence of all crown Asteroiidae synapomorphies,
796 as outlined here, challenges these classifications. *Savignasterias villieri* is known only from
797 isolated body wall ossicles, the shape of which leaves no doubt regarding its forcipulatid
798 affinities. Keyhole-shaped *psas* are present in *M. spiniger*, the Asteroiidae, the Stichasteridae and
799 the genus *Heliaster* but they are missing in *S. villieri* (Gale, 2011b). No adambulacral,
800 ambulacral or oral frame ossicles were described. The material remains too incomplete to assess
801 the species' phylogenetic position more clearly. Rousseau, Gale & Thuy (2018) discussed the
802 affinities of *P. janusensis* and concluded that even though no crossed pedicellariae have been
803 found and the unusual arrangement of the body wall skeleton for extant asteroiids, *P. janusensis*
804 should be considered an asteroiid because of its relatively short terminal ossicles, the strongly
805 quadriseriate arrangement of the tube feet, the morphology of its oral ossicles and basal piece of

Commenté [ST22]: Or not found yet !

Commenté [ST23]: You mean : diagnostic characters ?

806 straight pedicellariae. Although this character combination can be found in Asteroiidae, they are
807 also found in other forcipulatid groups such as the Stichasteridae or the polyphyletic family
808 Pedicellasteridae, and none of the listed characters were found to be synapomorphies of the
809 Asteroiidae here or by Fau & Villier (2020). A phylogenetic reappraisal of both *S. villieri* and *P.*
810 *janusensis* is still required but is beyond the scope of the current paper.

811 Blake & Guensburg (2016) reported a new fossil Stichasteridae from the Oxfordian Swift
812 Formation of Montana, *Atalopogaster gundersoni*. Unfortunately, the fossil's preservation does
813 not permit a detailed description. Blake & Guensburg, 2016 placed it within the family
814 Stichasteridae stating: "Based on overall shape, ossicular expression, and fusion of the arms,
815 *Atalopogaster* is aligned with *Neomorphaster* and the fossil genera *Argoviaster* Hess, 1972, and
816 *Pogaster* Blake and Peterson, 1993, in the Stichasteridae *sensu* Mah and Foltz (2011b)" (Blake &
817 Guensburg, 2016; p. 1161). Unfortunately, a more detailed investigation of the phylogenetic
818 position of *A. gundersoni* is not easily attempted owing to the poor preservation of the limited
819 number of specimens recovered to date.

820

821 **Cretaceous and younger forcipulataceans taxa**

822

823 Only six forcipulataceans species have been described from the Cretaceous to date,
824 among which three belong to the Zorocallida: *Protothyraea priscus* de Loriol, 1874, *Alkaidia*
825 *sumralli* Blake & Reid, 1998 and *Alkaidia megaungula* Ewin & Gale 2020 (Fau & Villier, 2023).
826 The three other forcipulataceans known to date are *Afraster scalariformis* Blake et al., 1996, an
827 assumed "pedicellasterid" from the Coniacian of Angola, the stichasterids *Pogaster stichos* Blake
828 & Peterson, 1993 from the Campanian of the USA, and the asteroiid *Cretasterias reticulatus* Gale

829 & Villier, 2013 from the Maastrichtian of Morocco. The oldest extinct taxa known to date to
830 present evidence of crossed pedicellariae arranged in wreath organs, an important synapomorphy
831 of the Asteriidae, is *C. reticulatus* (Gale & Villier, 2013). However, keyhole-shaped *psas* are
832 restricted to the actinals in *C. reticulatus* (Gale & Villier, 2013; personal observations), whereas
833 the presence of keyhole-shaped *psas* on the inferomarginals only, is one of the synapomorphies
834 of the extant Asteriidae. New observations of *Afraster scalariformis* have shown evidence for
835 preserved pedicellariae, both straight and crossed pedicellariae, and keyhole-shaped *psas* on
836 inferomarginals, questioning its systematic position among the family Pedicellasteridae (MF
837 personal observations). Investigation of the phylogenetic position of Cretaceous forcipulatacean
838 sea stars is therefore needed to better understand the origin and diversification of the modern
839 families.

840 Brisingids fossil record is almost nonexistent, with only one occurrence known to date,
841 *Hymenodiscus*, from the Miocene of Japan (Yamaoka, 1987). Mah & Foltz (2011a) argued upon
842 a late diversification of the Brisingida, because of their derived phylogenetic position and their
843 relatively young fossil record. Morphology-based phylogenies usually fail at recognizing a
844 derived position of the Brisingida within the Forcipulatacea (Gale 2011a; Fau & Villier, 2020). A
845 reappraisal of Cretaceous and Cenozoic forcipulataceans are therefore needed to understand the
846 complex evolutionary history of this group.

847

848 Conclusion

849

850 We taxonomically reevaluated six fossil taxa, placed these species in a phylogenetic
851 context, and constructed an evolutionary timeline for major diversification events in the history

852 of the Forcipulatacea. Our results provide substantial evidence for a delayed origination of the
 853 family Asteroiidae. The combined phylogenetic analysis of fossil and extant taxa suggests that the
 854 Jurassic forms exhibited unique combinations of characters that distinguish them from the extant
 855 families or genera they were previously assigned to and are characterized by intermediate
 856 morphological characters. This also implies a progressive acquisition of characters leading to the
 857 extant crown group families after the Jurassic. None of the eight Jurassic species analyzed here
 858 were placed within any of the extant families, but instead represent parts of stem-groups. The
 859 clade comprising the Asteroiidae and Stichasteridae has no fossil record before the Late
 860 Cretaceous. The absence of known Jurassic asteroiids suggest a Cretaceous or even younger origin
 861 for the clade, and phylogenetic divergence dating provides evidence in favor of a delayed
 862 origination and diversification of this major clade. Understanding the timing and pace of
 863 diversification of the Asteroiidae is of great interest to understanding recent biogeographical
 864 patterns and the undeniable evolutionary success of the family. Accumulation of the crown-
 865 group characters during the Mesozoic was more progressive than formerly accepted. Moreover,
 866 our results contradict the idea of a rapid diversification of the Forcipulatacea during the Triassic
 867 or the earliest Jurassic.

Commenté [ST24]: I would rather use "suggest" or write that there is no fossil evidence of an early evolution of the family : the fact that the analyzed fossils do not belong to the family does not mean that the family had not evolved yet (new fossils may be found in the future).

Commenté [ST25]: Intermediate between which forms ? Please precise.

a supprimé: do not

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Commenté [ST26]: Why ? Can you precise.

Commenté [ST27]: Do you mean in terms of species richness ? Please precise.

Commenté [ST28]: You mean Evolution ?

Commenté [ST29]: Or at least : "do not support"

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879

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891 *mêmes, d'après l'état actuel de nos connaissances, soit relativement à l'utilité qu'en*
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